The background of the slide is a photograph of a sunset or sunrise. The sky is a gradient of colors from light blue at the top to orange and red near the horizon. There are wispy clouds and a few bright spots, possibly stars or distant lights. In the foreground, the dark silhouettes of two large trees are visible against the bright sky.

Explicit simulation of gravity waves up to the lower thermosphere using an idealized GCM

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Why a GW resolving GCM?

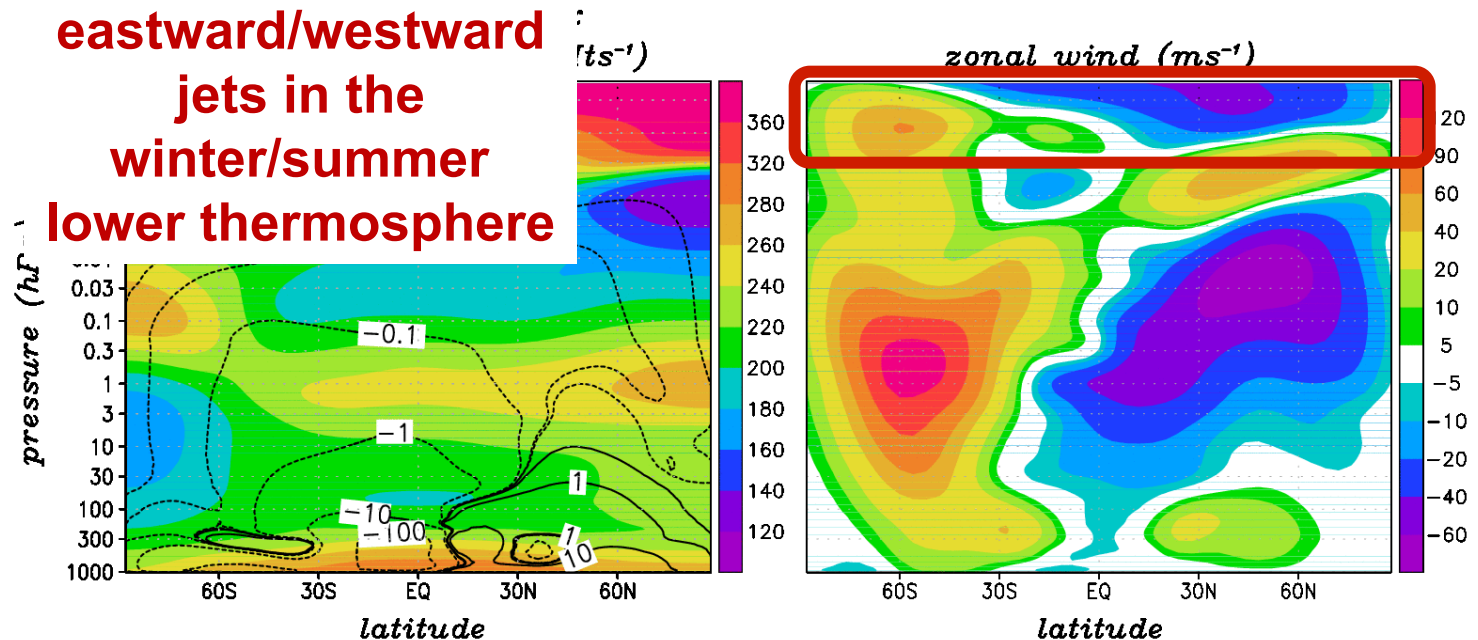
Current GW parameterizations do a good job in providing the GW drag that is required in the mesosphere (non-orographic GWs) and winter stratosphere (orographic GWs), but **are subject to the strong assumptions of single column dynamics and a quasi-stationary GW energy equation.**

- Interaction with resolved waves and parameterized turbulent diffusion is generally questionable.
- Continuous sources for *non-orographic waves* cannot be specified, only *launch-level parameters*.
- Intermittency is difficult to include.
- Parameterized GWs do not participate in the horizontal energy cascade (and in the Lorenz energy cycle).

Kühlungsborn Mechanistic general Circulation Model (KMCM)

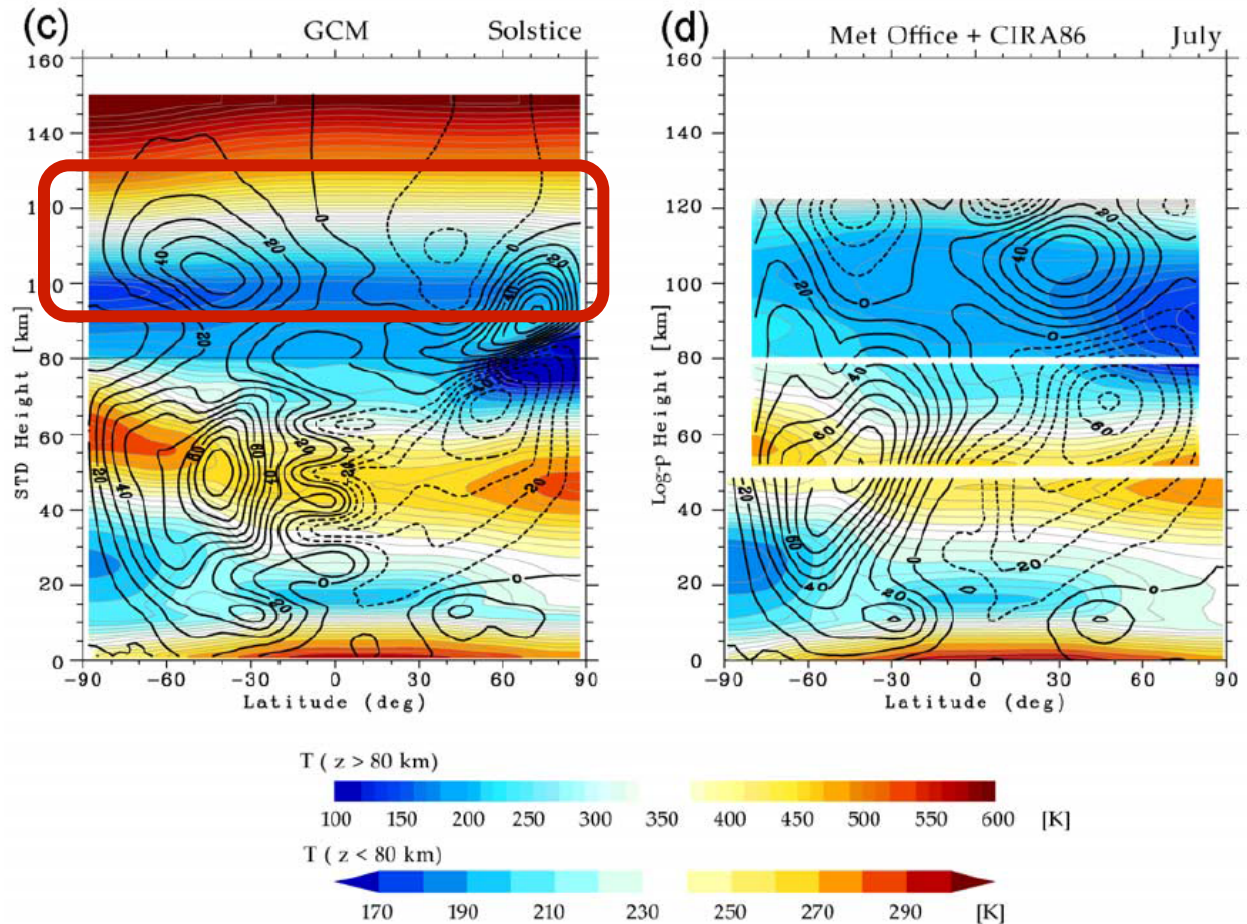
- hydrostatic spectral GCM (Simmons & Burridge, 1981, MWR); T120L190 or T240L190 (up to ~130 km, level spacing ~600 m up to 100 km)
- tracer transport based on the spectral method (Schlutow et al., 2014, JGR)
- radiative transfer and tropospheric moisture cycle, including a slab ocean, land-sea masks (Becker et al., 2015, JASTP); vertically varying heat capacity and gas constant, molecular diffusion, and ion drag
- Smagorinsky-type horizontal and vertical diffusion; both diffusion coefficients depend on the Richardson number; additional linear horizontal diffusion in the lower thermosphere

July
climatology

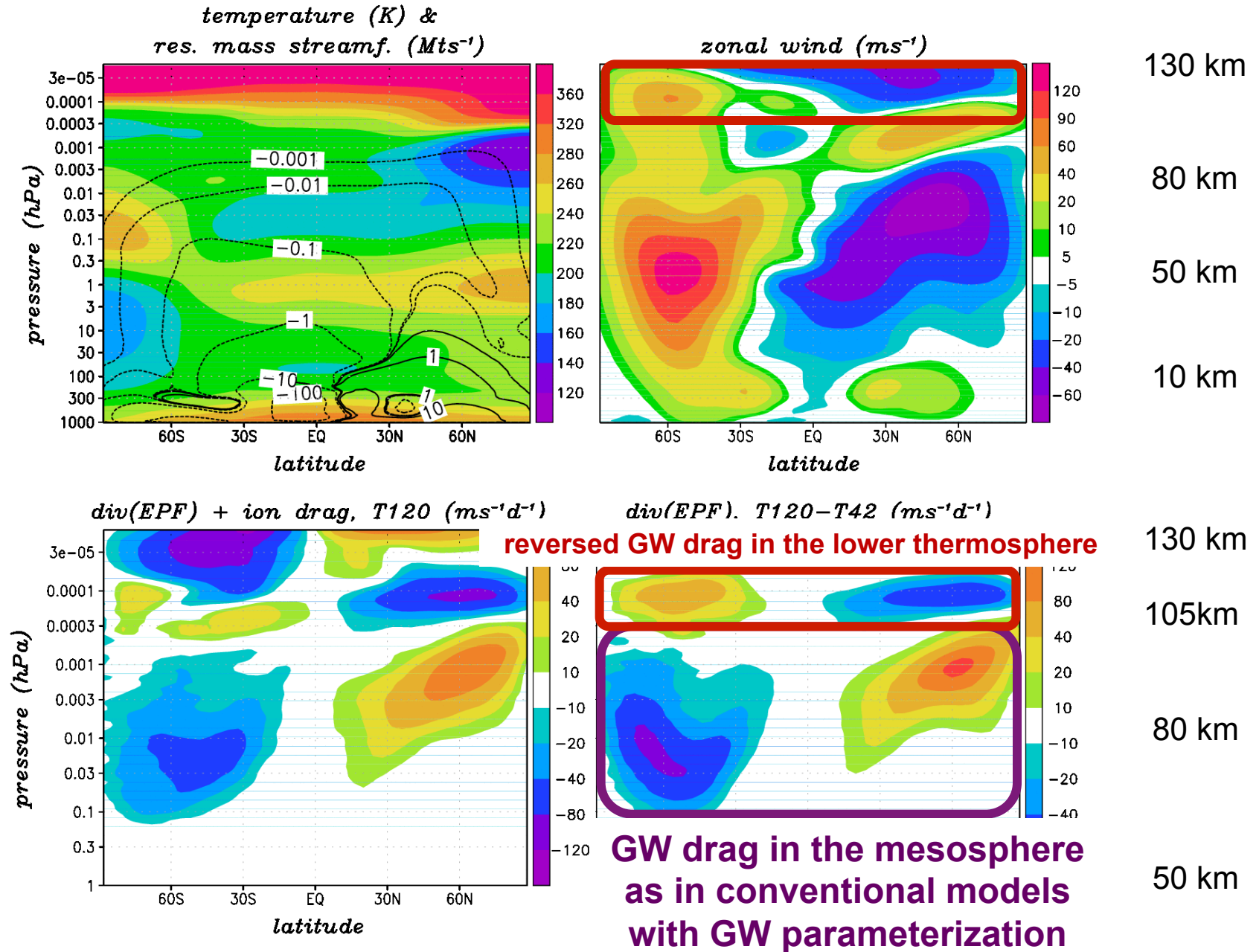


Another GW-resolving GCM with a realistic summer-to-winter-circulation in the upper mesosphere is the vertically extended KANTO model (Watanabe et al., 2008, JGR) called JAGUAR (Watanabe and Miyahara, 2009, JGR) which is a fully comprehensive spectral GCM with T213L270 resolution.

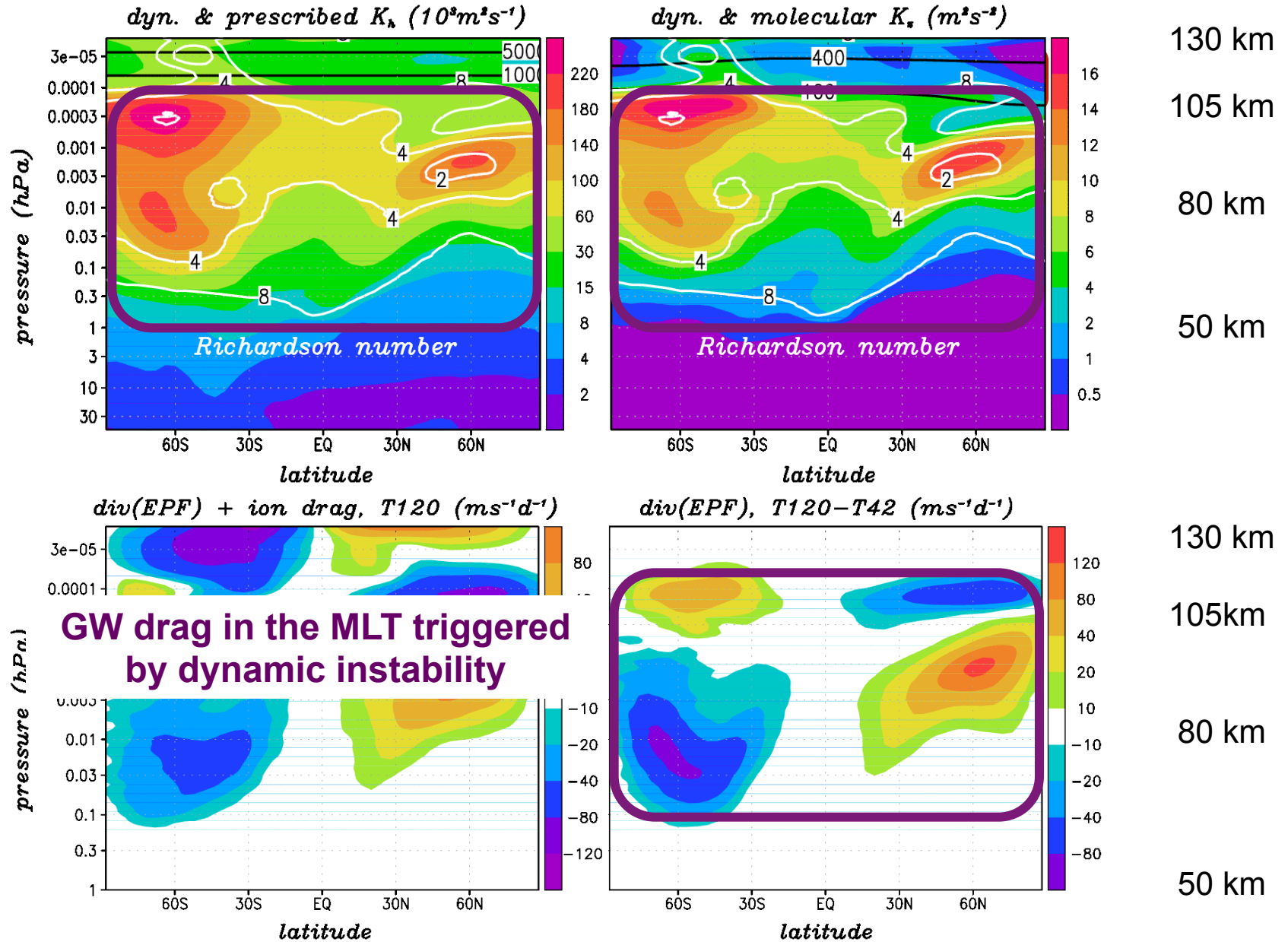
**eastward/westward
jets in the
winter/summer
lower thermosphere**



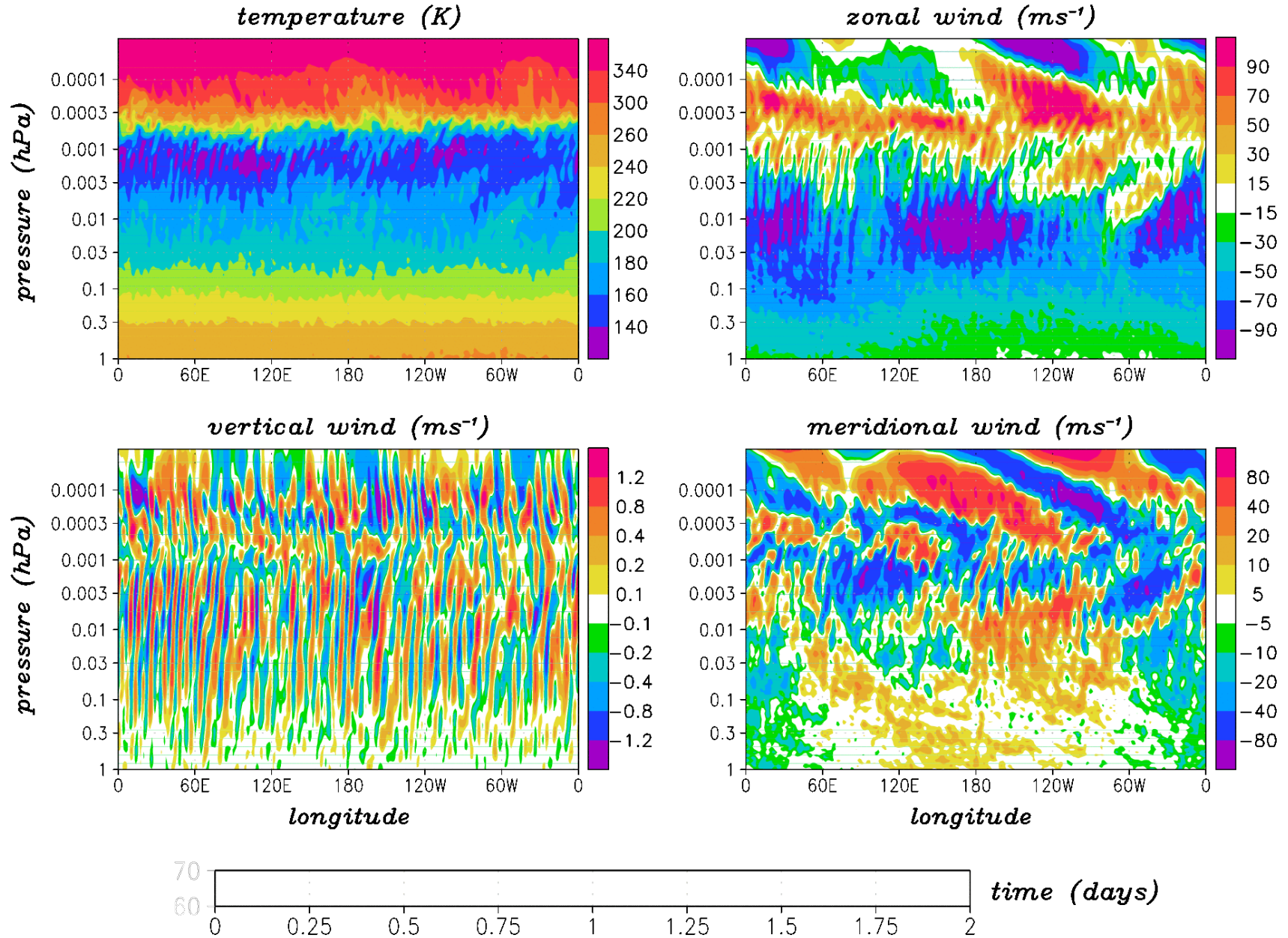
Zonal-mean climatology and wave driving, July, KMCM T120L190



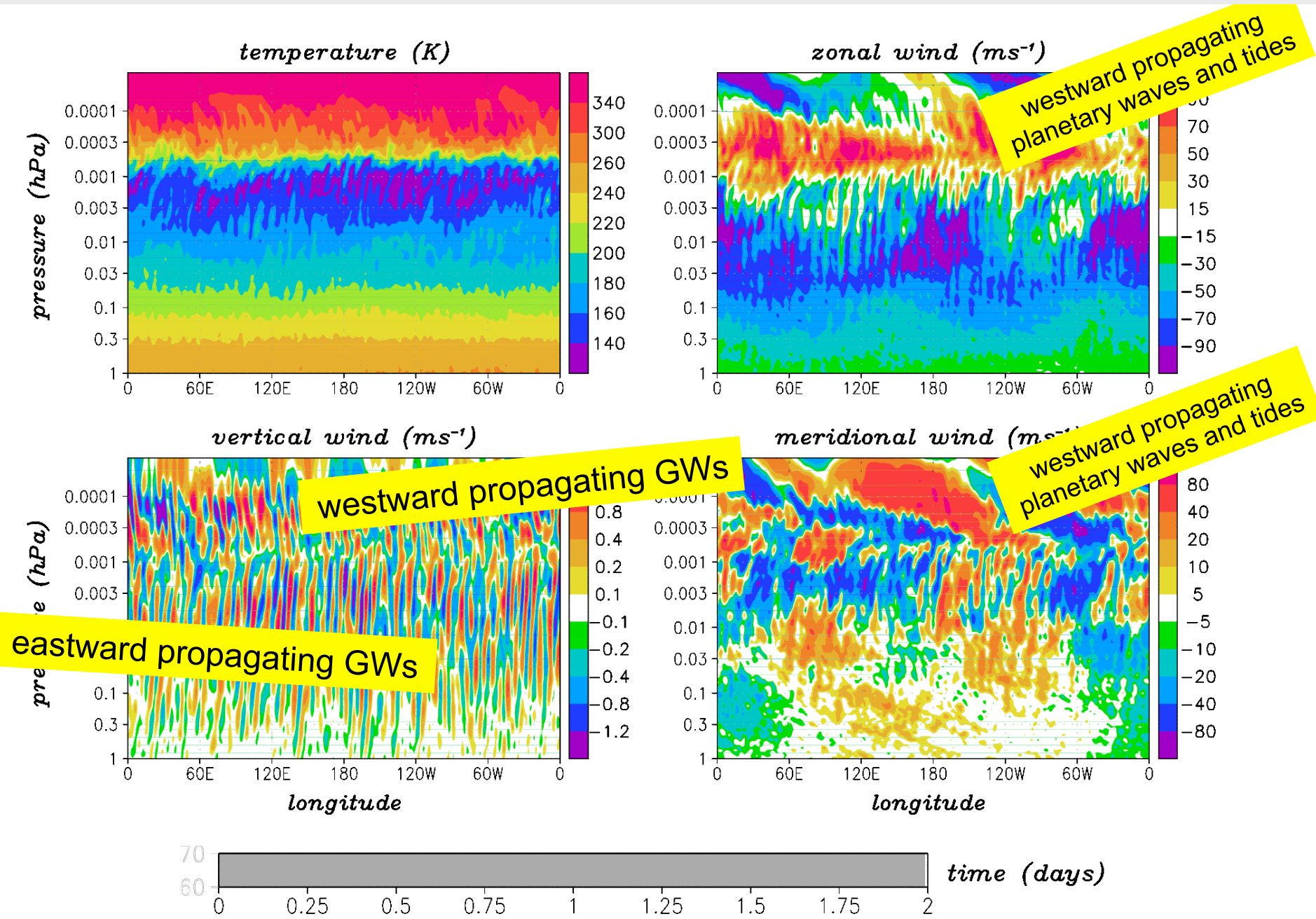
Diffusion scheme and wave driving, July, KMCM T120L190



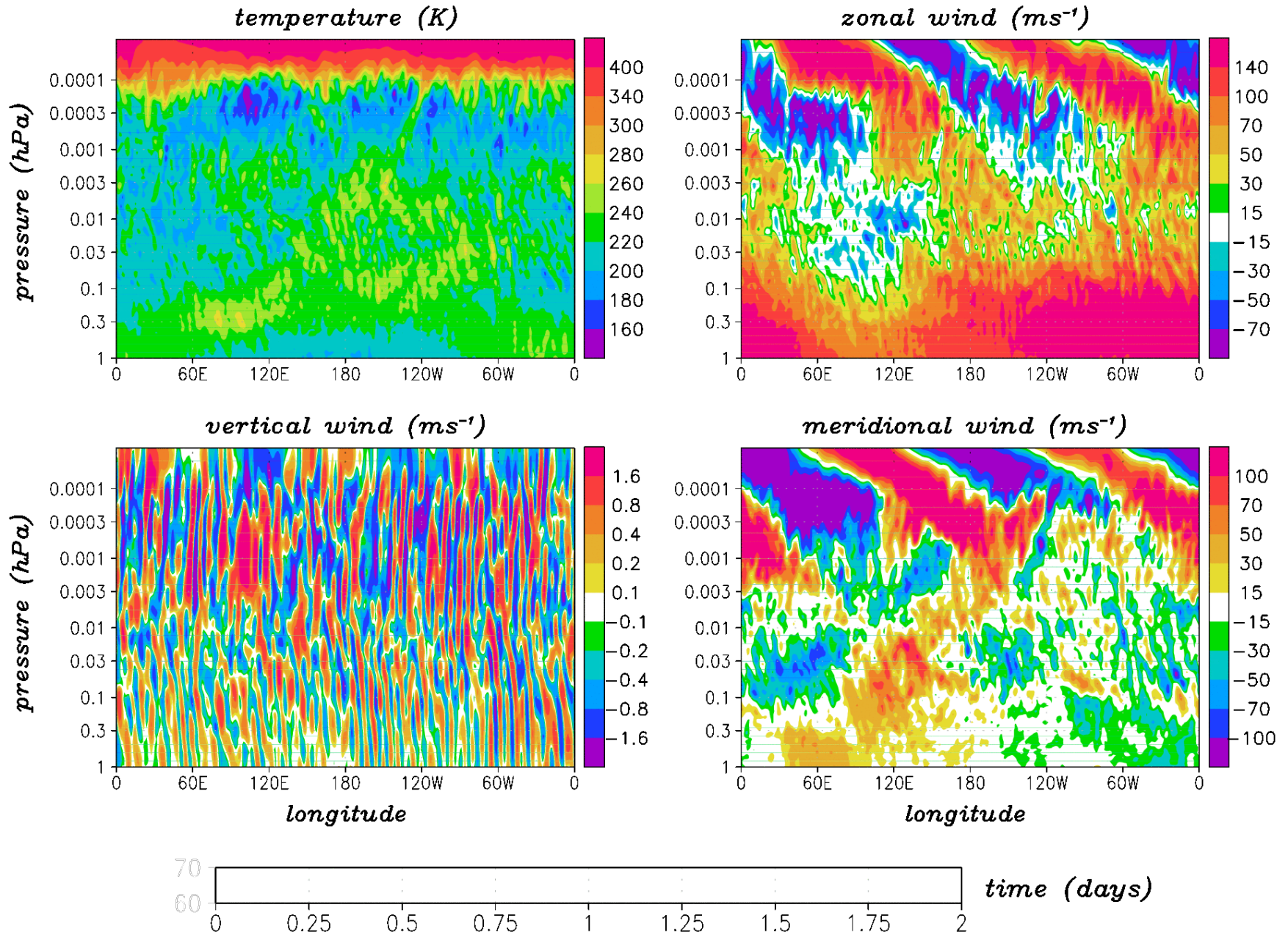
Variability in the MLT at 55°N during July



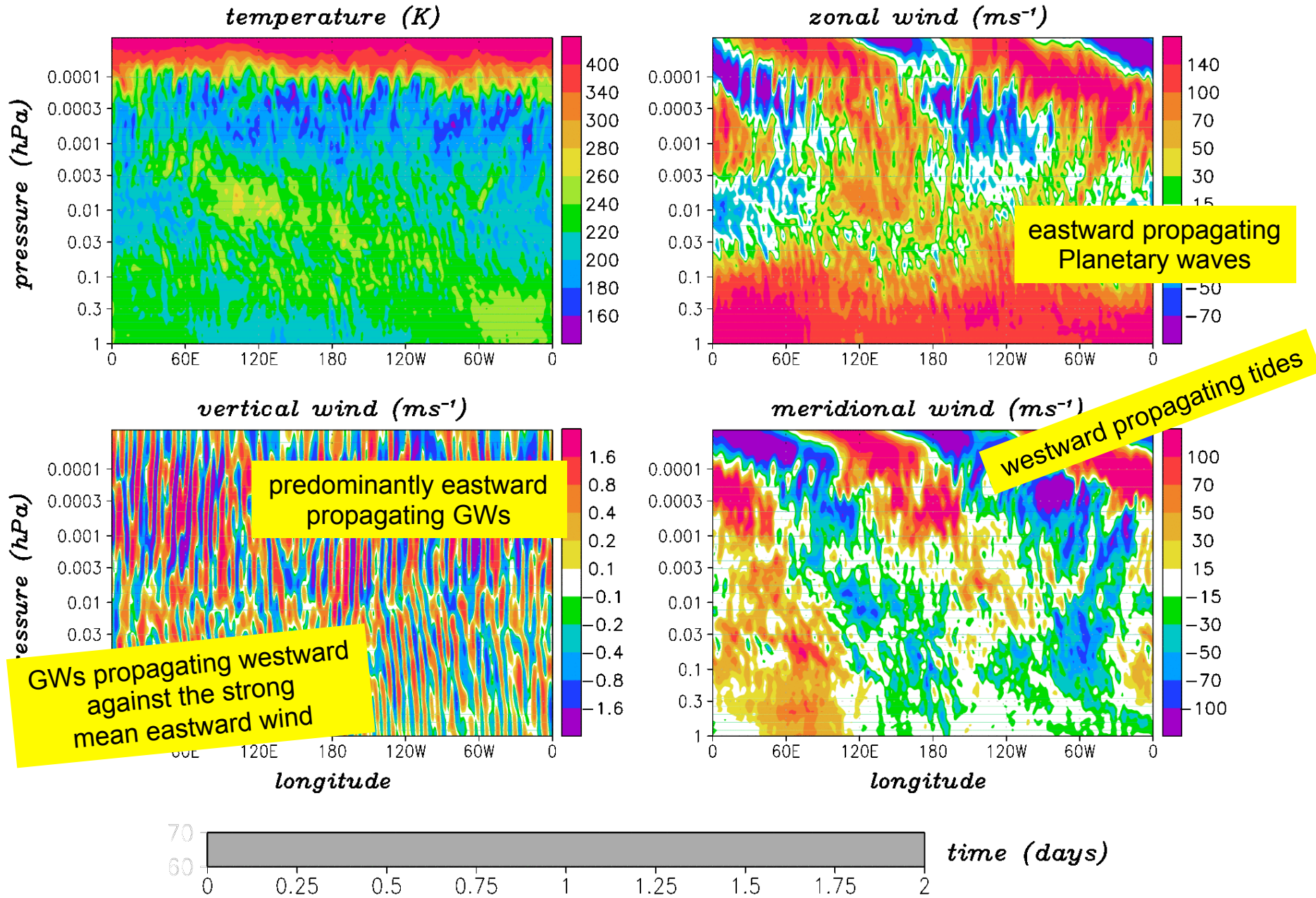
Variability in the MLT at 55°N during July



Variability in the MLT at 55°S during July



Variability in the MLT at 55°S during July



Simple wavenumber-frequency analysis to explore the nature of the reversed GW momentum flux in the lower thermosphere: Secondary GWs?

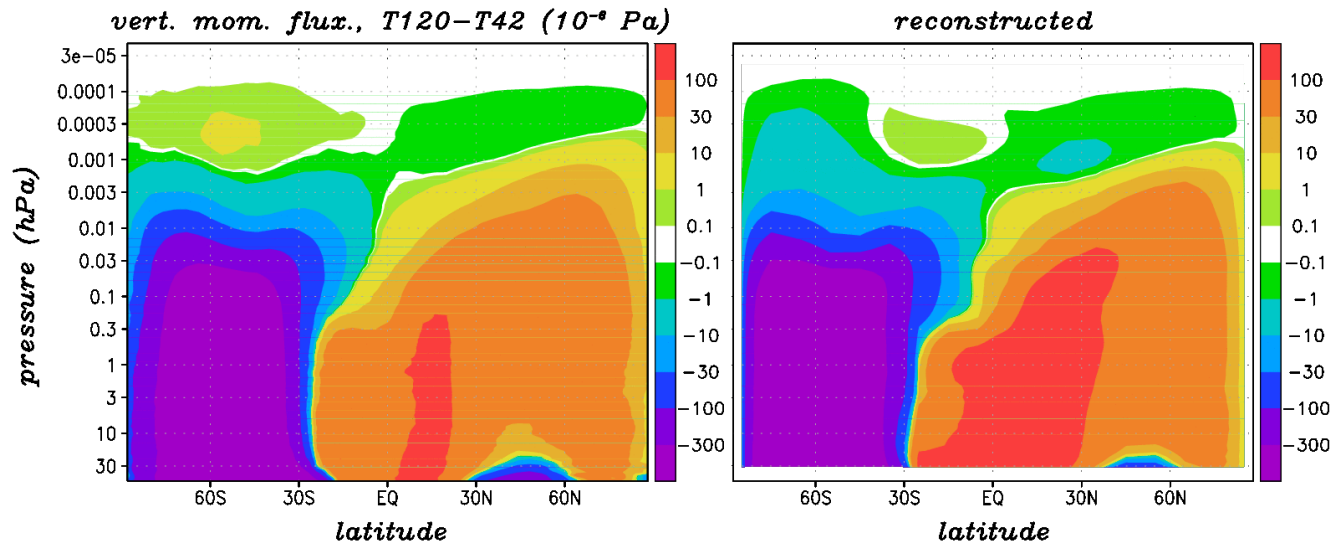
Assumptions: Zonal propagation, mid-frequency approximation
(It's better to be approximately right than precisely wrong)

$$u' = \text{Re} \left(\sum_m \sum_j \hat{u}_{mj}(\phi, p) e^{i(m\lambda - \omega_j t)} \right)$$

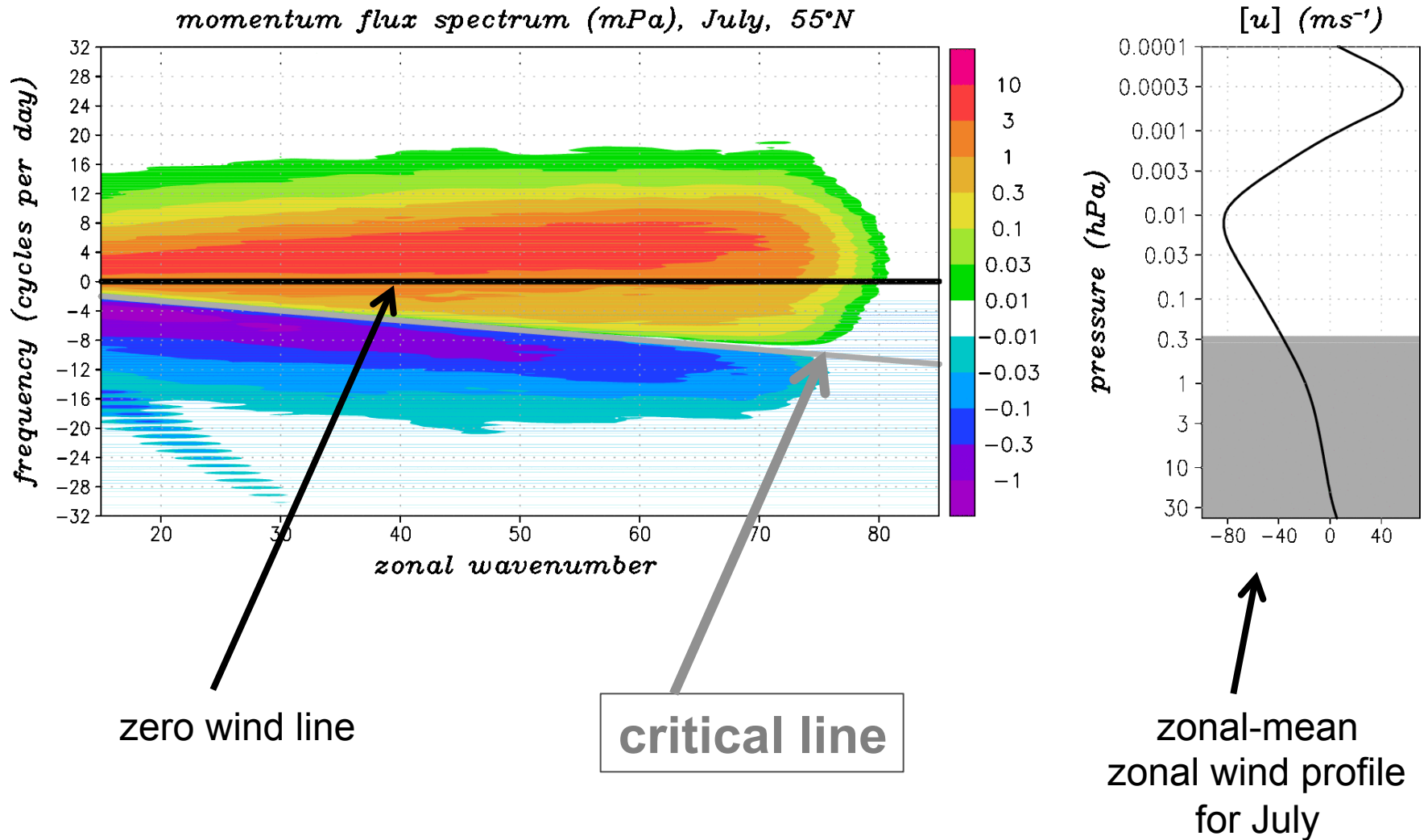
$$w' = -\text{Re} \left(\frac{m}{a \cos \phi} \frac{U - c_{mj}(\phi)}{N} \sum_m \sum_j \hat{u}_{mj}(\phi, p) e^{i(m\lambda - \omega_j t)} \right)$$

$$c_{mj}(\phi) = \omega_j a \cos \phi / m$$

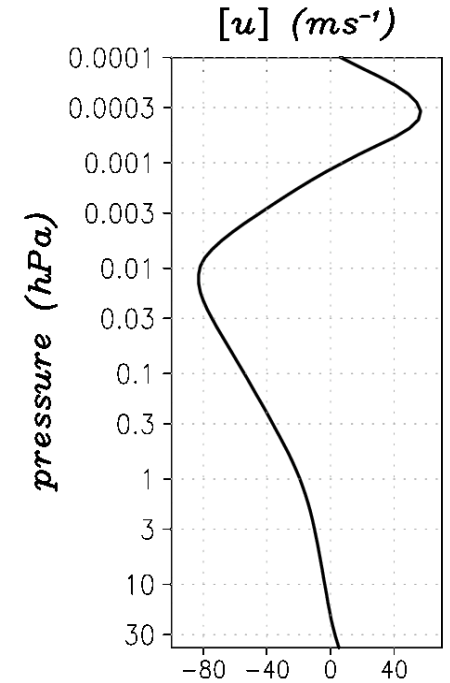
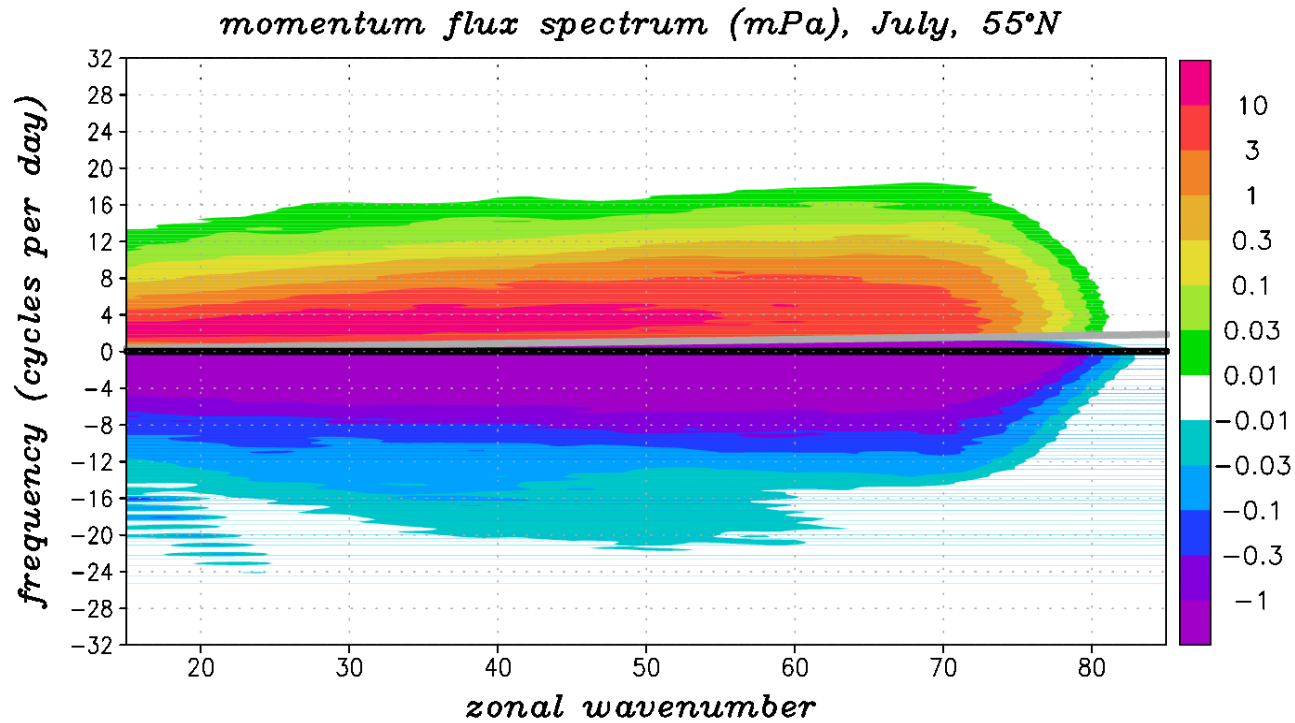
$$\rho \overline{u'w'} = -\frac{\rho}{2} \sum_m \sum_j \frac{m}{a \cos \phi} \frac{U - c_{mj}(\phi)}{N} |\hat{u}_{mj}(\phi, p)|^2$$



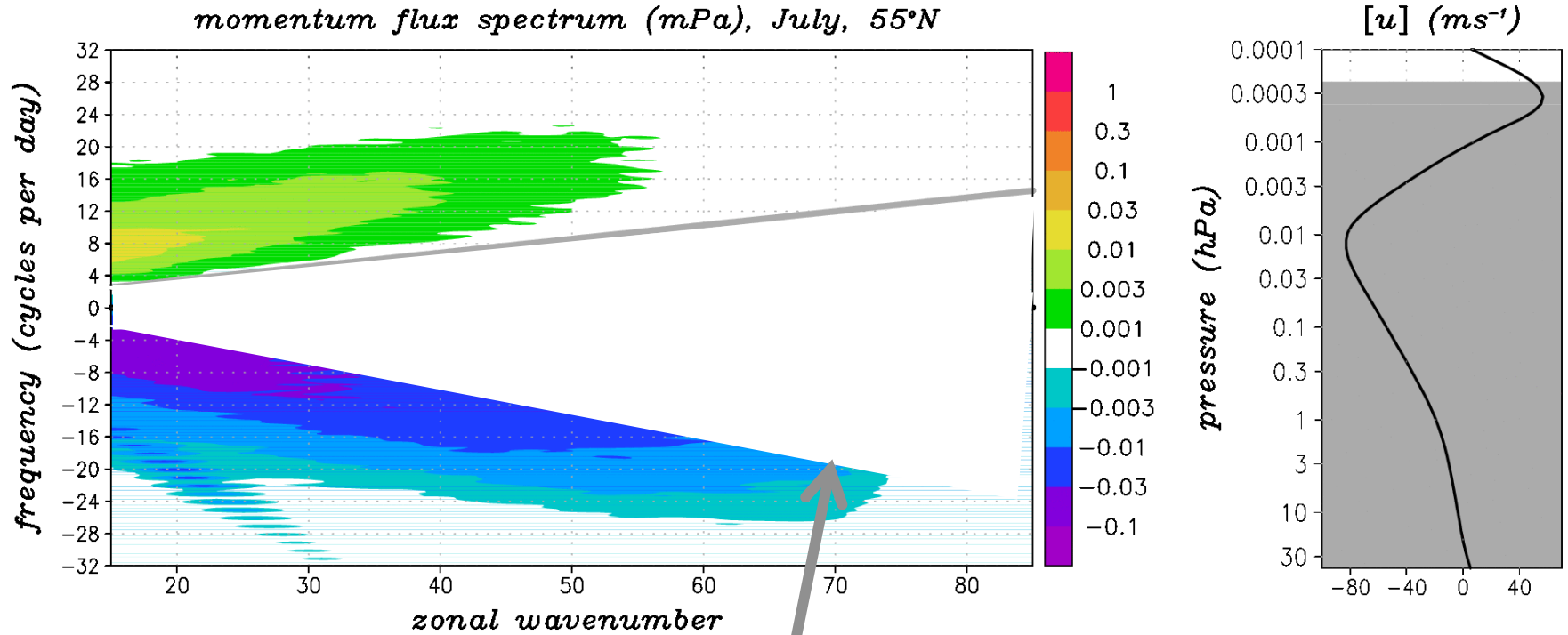
Vertical evolution of the GW momentum flux spectrum (July, 55°N)



Vertical evolution of the GW momentum flux spectrum (July, 55°N)

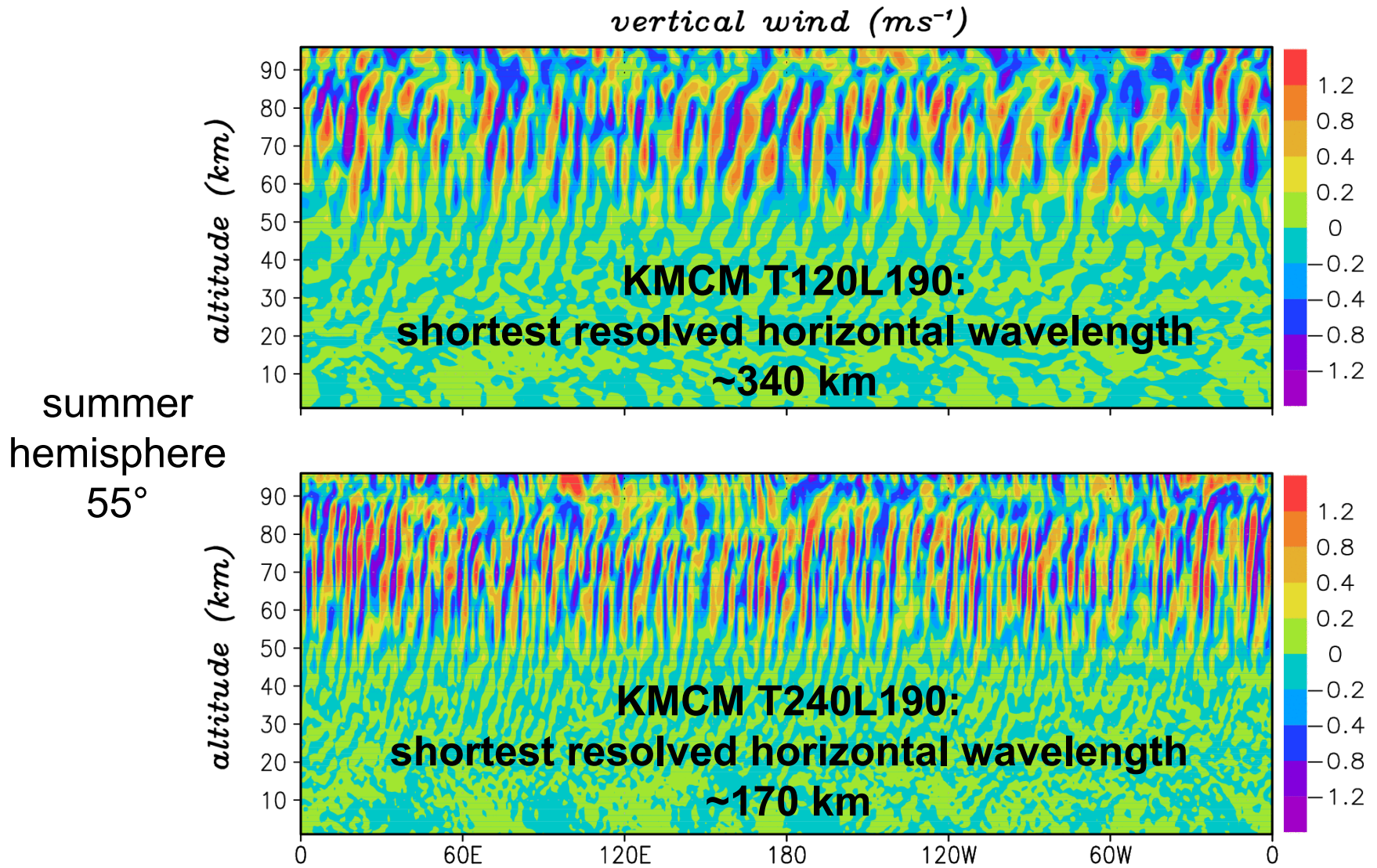


Vertical evolution of the GW momentum flux spectrum (July, 55°N)



**GW modes
supposed
to be absorbed
at critical levels**

Resolution dependence of the simulated GW scales (former model version)<

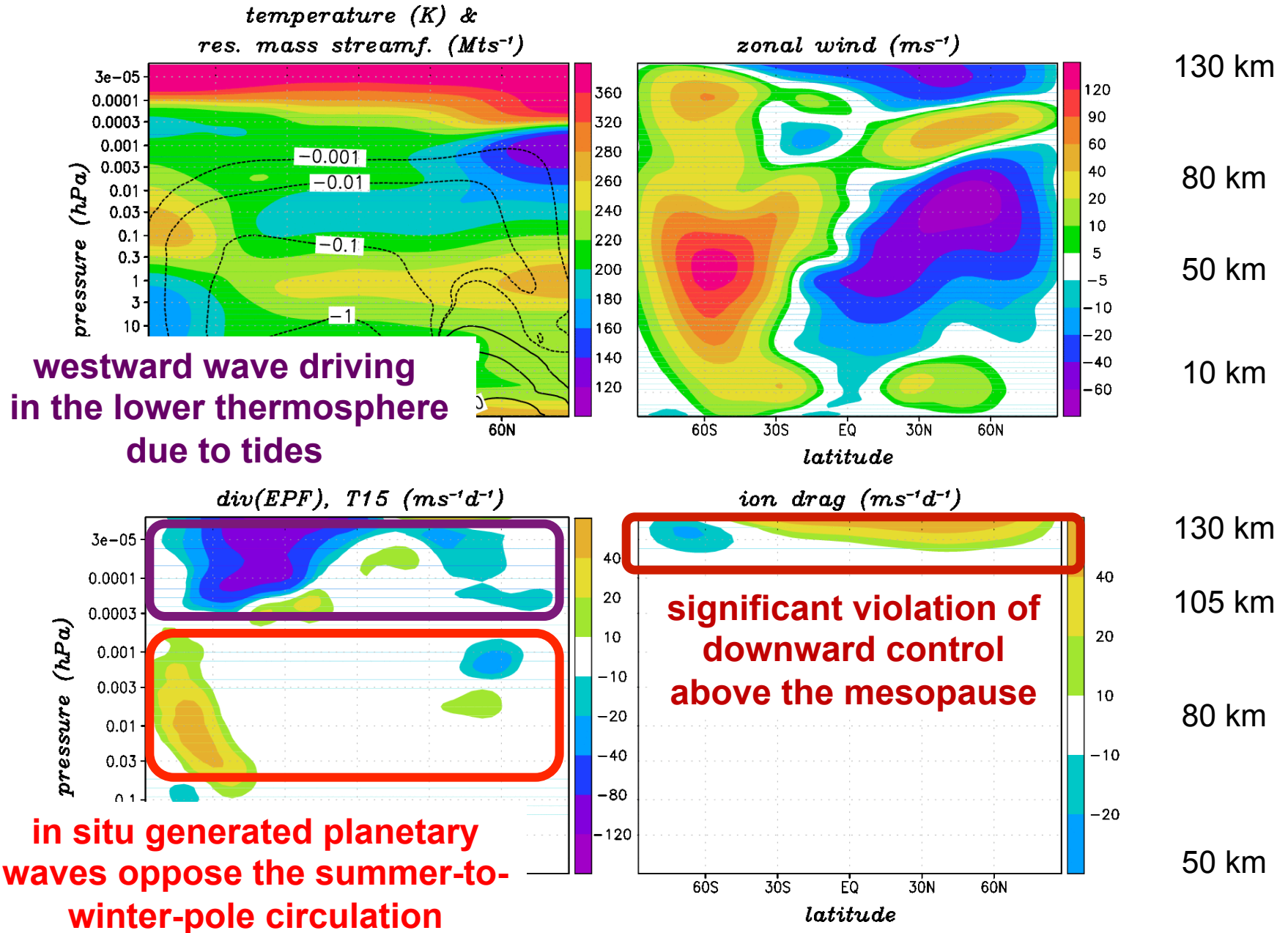


For higher resolution, the simulated gravity waves have smaller spatial scales (and higher frequencies).

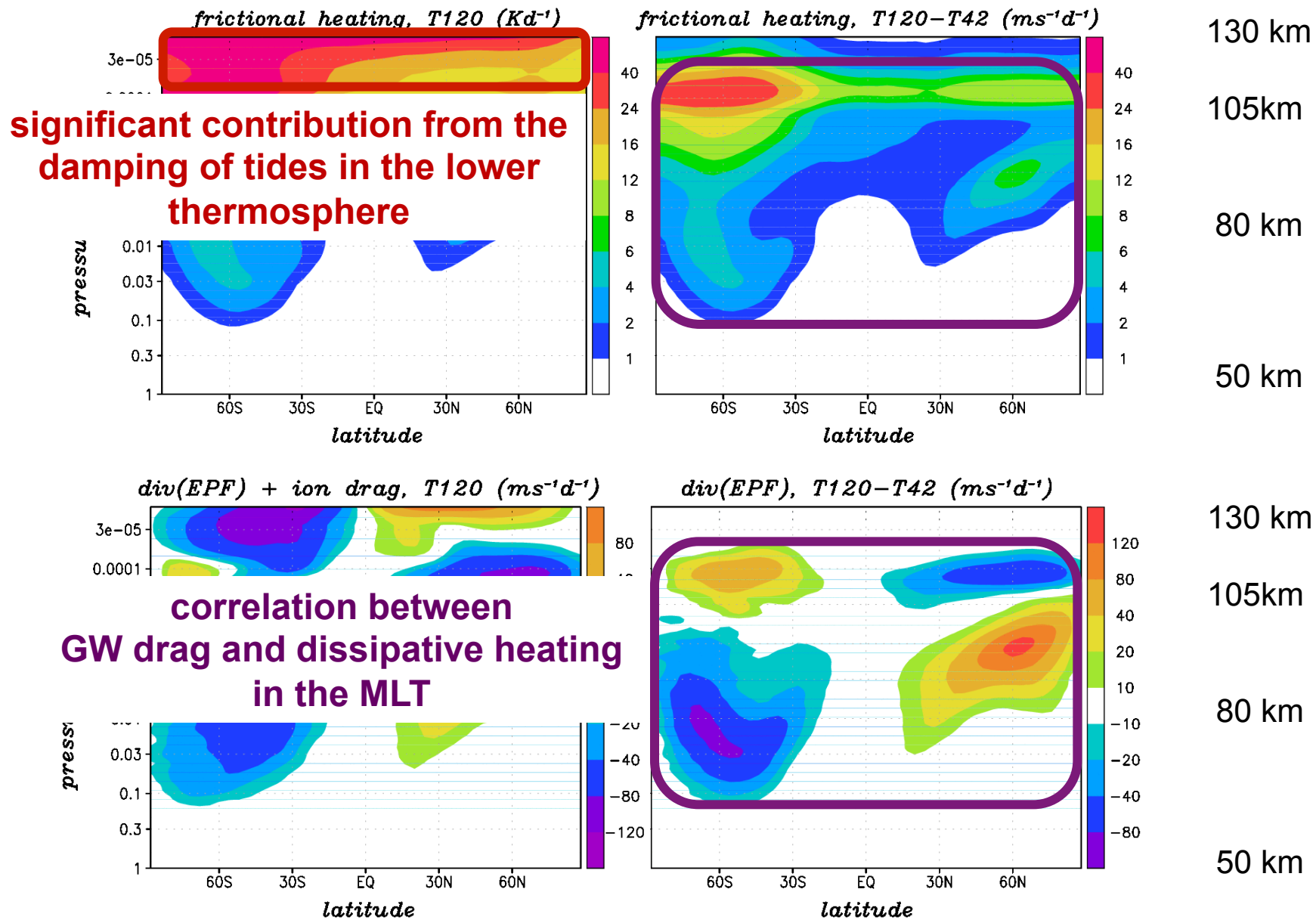
Conclusions

- A GW-resolving (spectral) GCM with a Smagorinsky-type diffusion scheme that adjusts to dynamic instability (for vertical and horizontal diffusion) yields realistic non-orographic GW drag in the MLT.
- **GW attenuation at critical levels at lower altitudes where the GW amplitudes are small is not resolved for a vertical level spacing of a few 100 m.** The simulated reversal of the momentum flux above the mesopause is spurious.
- The simulated GW scales are resolution-dependent (general issue of non-convergence of circulation models with respect to numerical resolution).

Zonal-mean climatology and wave driving, July, KMCM T120L190

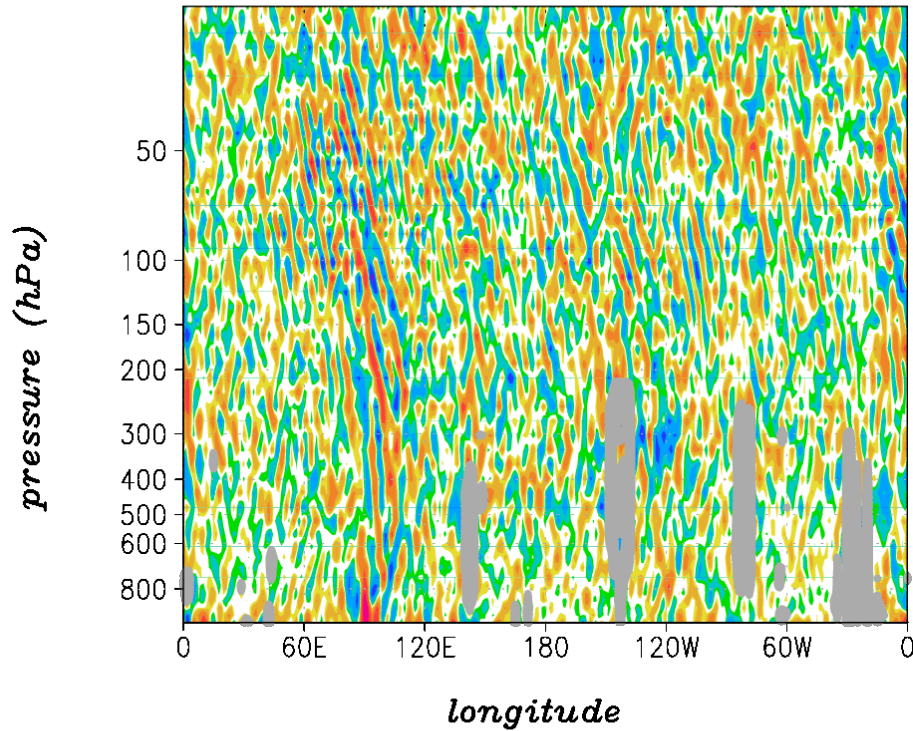


Frictional heating and wave driving, July, KMCM T120L190

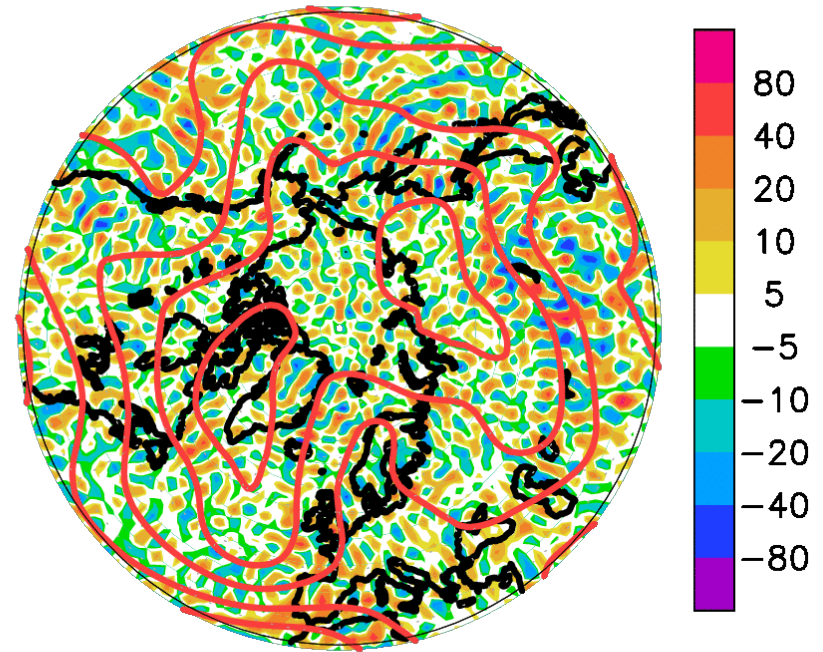


Tropospheric GW activity: Arbitrary 2-day time interval during July

horiz. divergence (10^{-6}s^{-1} , 45°N)
latent heating $> 2 \text{ Kd}^{-1}$



horiz. divergence (10^{-6}s^{-1}), 200 hPa
horiz. streamf.



Subgrid-scale diffusion:
 Classical anisotropic Smagorinsky Model
 extended by a Richardson criterion (Becker, 2009, JAS)

$$(\partial_t \mathbf{v})_{diff} = \rho^{-1} \nabla_h (\rho K_h \mathbf{S}_h) + \rho^{-1} \partial_z (\rho K_z \partial_z \mathbf{v})$$

$$\mathbf{S}_h = \{ (\nabla + \mathbf{e}_z/a) \circ \mathbf{v} \} + \{ (\nabla + \mathbf{e}_z/a) \circ \mathbf{v} \}^T - \mathbf{1} \nabla \cdot \mathbf{v}$$

$$(\partial_t c_p T)_{diff} = \frac{c_p}{\rho} \nabla_h (\rho K_h \nabla_h T) + \frac{c_p}{\rho} \partial_z \left(\rho \frac{T}{\Theta} K_z \partial_z \Theta \right) \\ + K_h |\mathbf{S}_h|^2 + K_z (\partial_z \mathbf{v})^2$$

$$K_h = \left(l_h^2 \right) |\mathbf{S}_h| \left(1 + \left(\alpha \right) F(R_i) \right) \gg K_z = \left(l_z^2 \right) |\partial_z \mathbf{v}| F(R_i)$$

$$F(R_i) = \begin{cases} \sqrt{1 - 18 R_i} & R_i < 0 \\ (1 + 9 R_i)^{-1}, & R_i \geq 0 \end{cases}$$

$$R_i = N^2 / (\partial_z \mathbf{v})^2 \rightarrow R_i - \frac{1}{4} \quad \text{for } p < 100 \text{ hPa}$$

3 free parameters: 2 mixing lengths and α
(Prandtl numbers are assumed to be 1)

Dynamical Smagorinsky Model

(Schaefer-Rolffs et al-, 2014, Met. Z.)

$$(\partial_t \mathbf{v})_{diff} = \rho^{-1} \nabla_h (\rho K_h \mathbf{S}_h) + \rho^{-1} \partial_z (\rho K_z \partial_z \mathbf{v})$$

$$\mathbf{S}_h = \{(\nabla + \mathbf{e}_z/a) \circ \mathbf{v}\} + \{(\nabla + \mathbf{e}_z/a) \circ \mathbf{v}\}^T - \mathbf{1} \nabla \cdot \mathbf{v}$$

$$(\partial_t c_p T)_{diff} = \frac{c_p}{\rho} \nabla_h \left(\rho \frac{T}{\Theta} K_h \Theta \nabla_h \Theta \right) + \frac{c_p}{\rho} \partial_z \left(\rho \frac{T}{\Theta} K_z \Theta \partial_z \Theta \right) \\ + K_h |\mathbf{S}_h|^2 + K_z (\partial_z \mathbf{v})^2$$

$$K_h = l_h^2 |\mathbf{S}_h| \gg K_z = l_z^2 |\partial_z \mathbf{v}|$$

\tilde{X} = average of X over the smallest resolved scales

$$l_h^2 = \left| \widetilde{\mathbf{v} \circ \mathbf{v}} - \tilde{\mathbf{v}} \circ \tilde{\mathbf{v}} - \frac{1}{2} \mathbf{1} (\widetilde{\mathbf{v}^2} - \tilde{\mathbf{v}}^2) \right| \left| \left(\frac{\tilde{\Delta}}{\Delta} \right)^2 |\widetilde{\mathbf{S}_h} \widetilde{\mathbf{S}_h} - |\mathbf{S}_h| \mathbf{S}_h \right|^{-1}$$

$$l_z \propto l_h^{1/3}$$

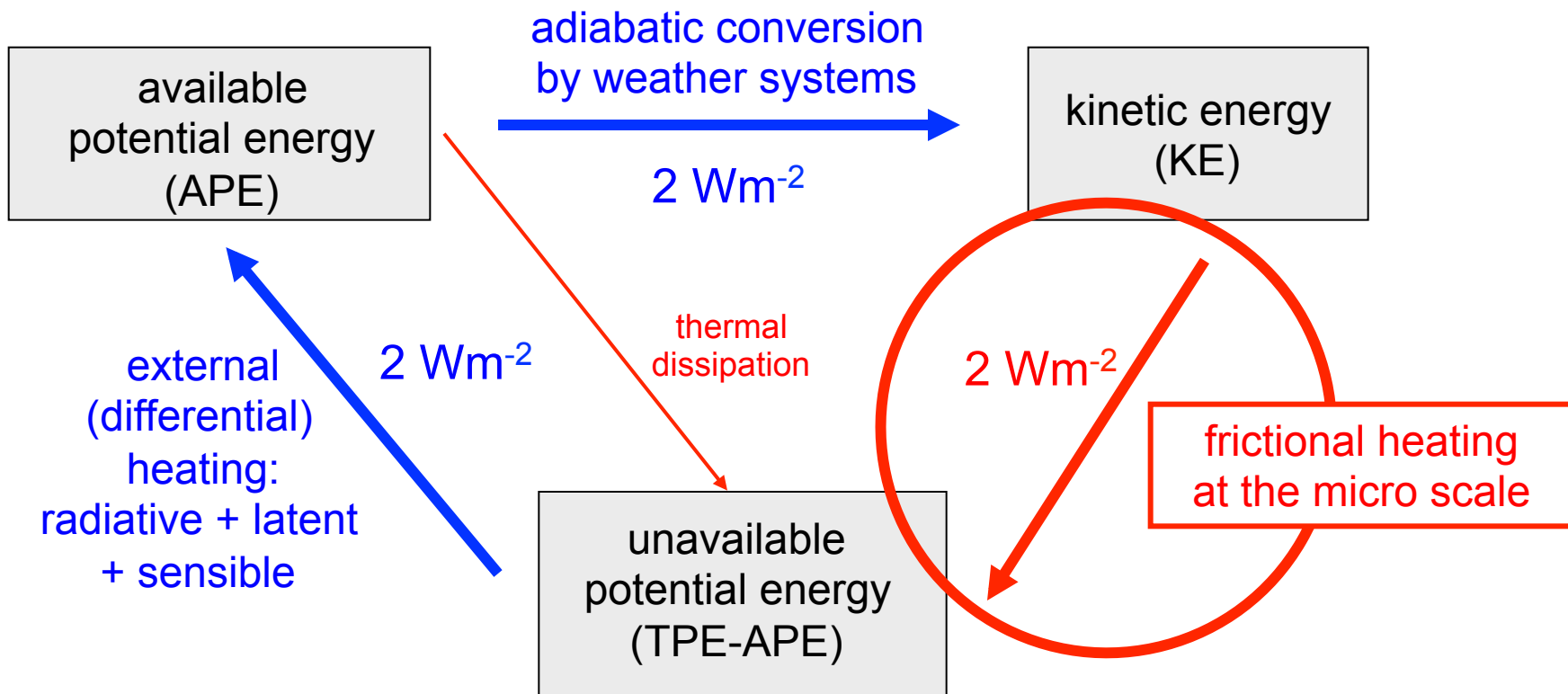
$$K_h \Theta = l_{h\Theta}^2 |\mathbf{S}_h| \gg K_z \Theta = l_{z\Theta}^2 |\partial_z \mathbf{v}|$$

**free parameters:
grid scale, filter scale, ...**

$$l_{h\Theta}^2 = \left| \widetilde{\Theta \mathbf{v}} - \tilde{\Theta} \tilde{\mathbf{v}} \right| \left| \left(\frac{\tilde{\Delta}}{\Delta} \right)^2 |\widetilde{\mathbf{S}_h} \widetilde{\nabla \Theta} - |\mathbf{S}_h| \widetilde{\nabla \Theta} \right|^{-1}$$

$$l_{z\Theta} \propto l_{h\Theta}^{1/3}$$

Lorenz energy cycle



- frictional heating essential for entropy and energy budgets: net diabatic heating of the atmosphere = frictional heating (Lorenz, 1967)
- frictional heating occurs at the end of energy cascades through the mesoscales (including GWs) and turbulence
- relative importance of thermal dissipation?

Energy transfers in conventional atmospheric models

incomplete
/ flawed
dissipation
of KE
and APE

TPE-APE (resolved flow)
(wavelengths larger than 500 km x 3 km)

differential heating

APE (resolved flow)
(wavelengths larger than 500 km x 3 km)

dynamics

KE (resolved flow)
(wavelengths larger than 500 km x 3 km)

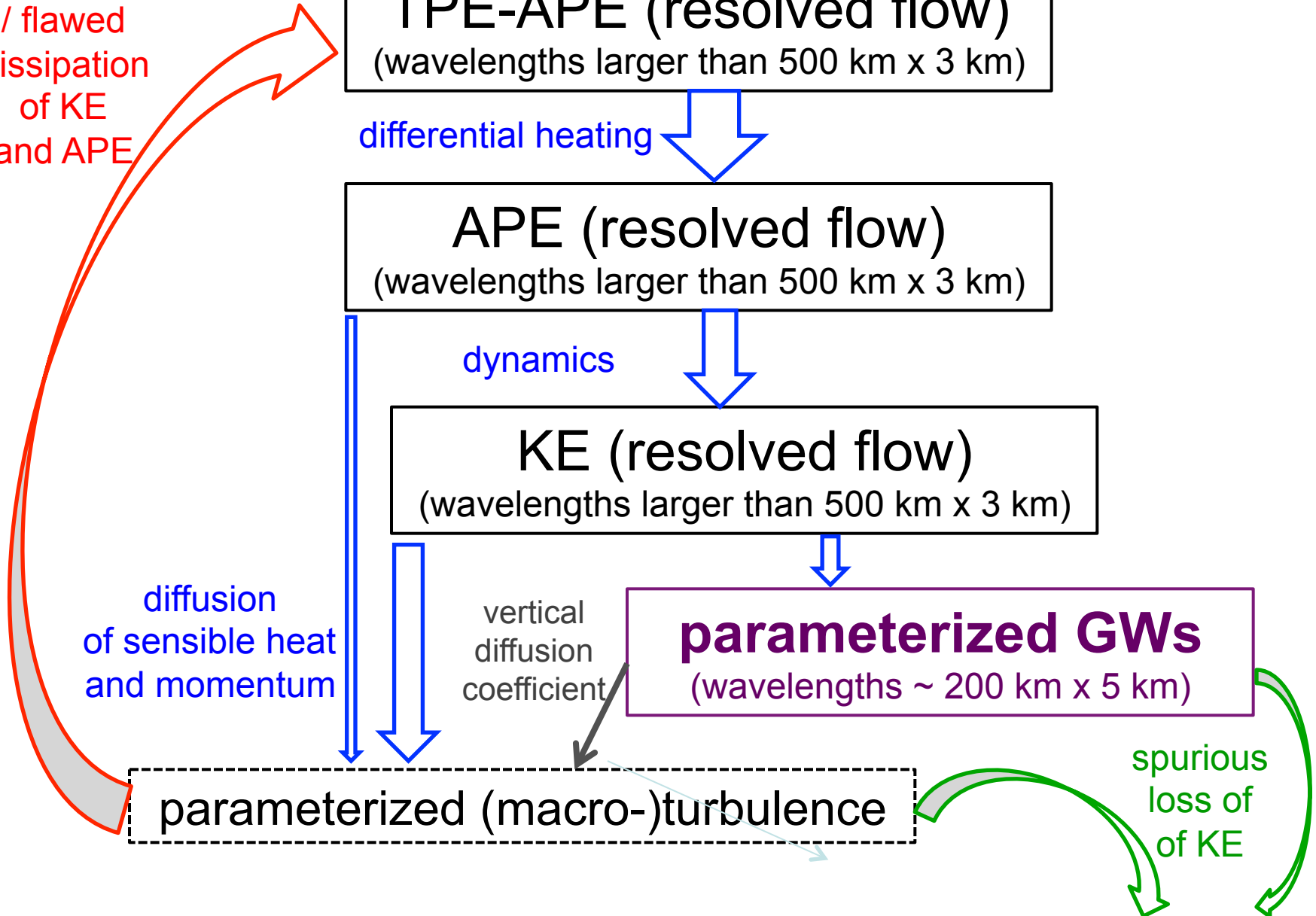
diffusion
of sensible heat
and momentum

vertical
diffusion
coefficient

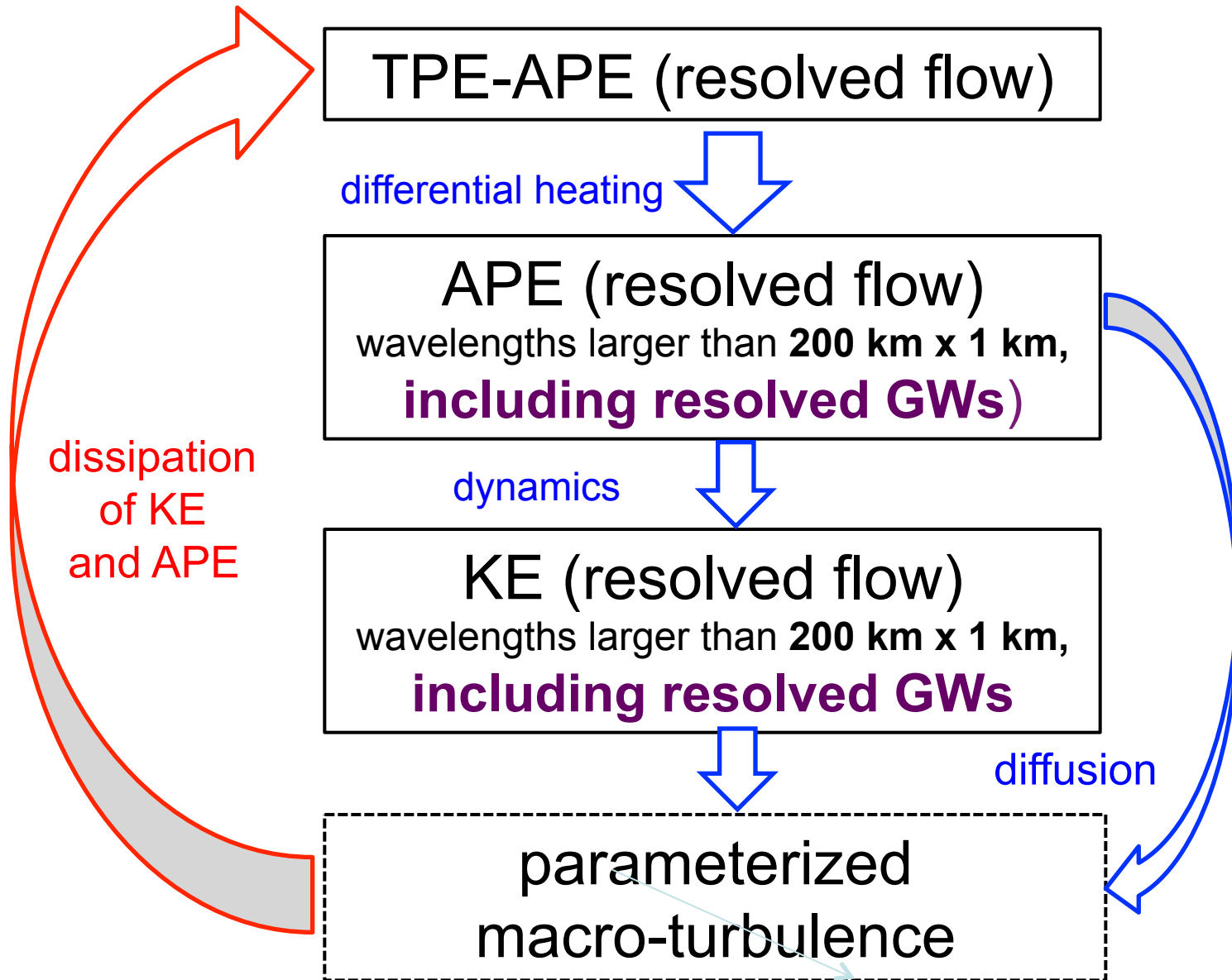
parameterized GWs
(wavelengths ~ 200 km x 5 km)

parameterized (macro-)turbulence

spurious
loss of
of KE



Energy transfers in a GW-resolving GCM with consistent subgrid-scale diffusion

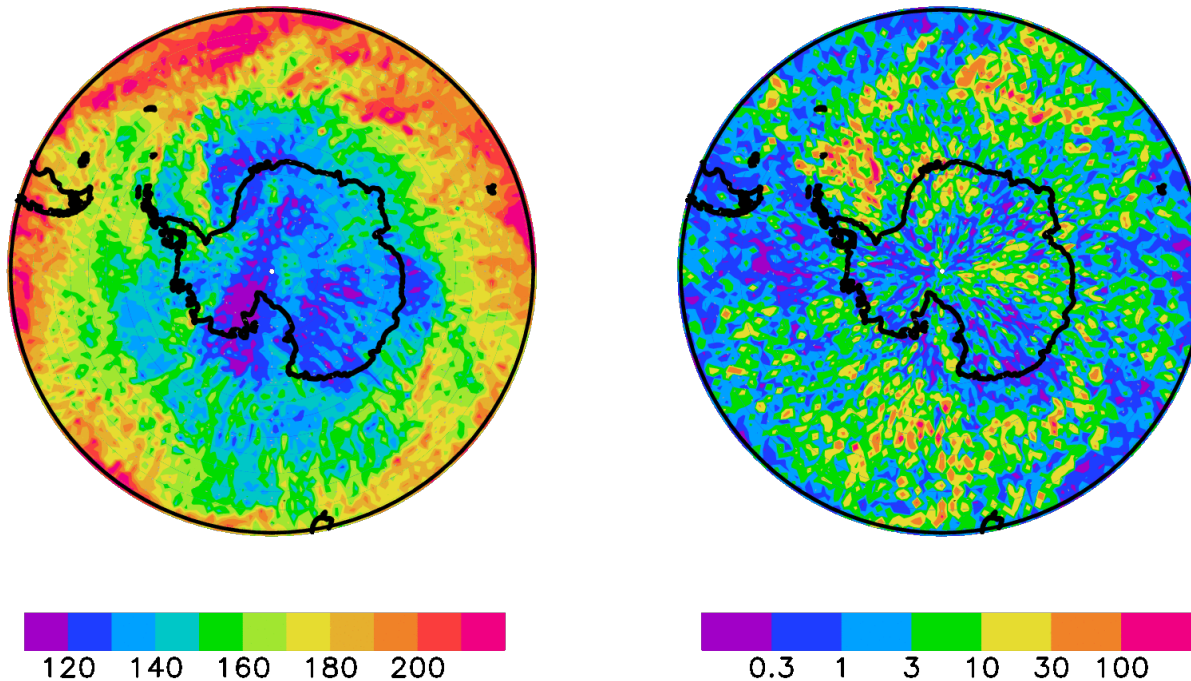


Why a gravity-wave resolving GCM?

GW parameterizations miss elementary aspects of the mesoscales in the real atmosphere:

- **Wave effects are highly intermittent (temporally and spatially)**

snapshot of temperature (K) and dissipation (K/d) around 85 km (January)



- **Parameterized GWs do not participate in the horizontal energy cascade, nor do they participate in the Lorenz energy cycle (except for the surface friction due to orographic GWs).**